

Physics Higher level Paper 2

30 April 2025

Zone A morning | Zone B morning | Zone C morning

C	Candida	ate se	ssion	numb	er

2 hours 30 minutes

Instructions to candidates

- Write your session number in the boxes above.
- Do not open this examination paper until instructed to do so.
- · Answer all questions.
- Answers must be written within the answer boxes provided.
- · A calculator is required for this paper.
- A clean copy of the **physics data booklet** is required for this paper.
- The maximum mark for this examination paper is [90 marks].

2225-6715

© International Baccalaureate Organization 2025

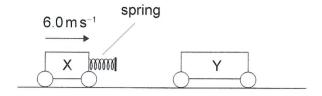




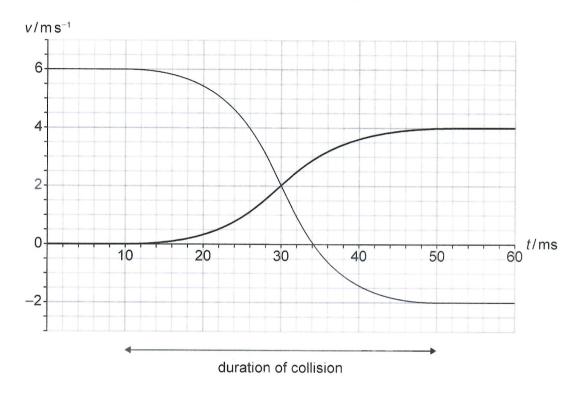
24 pages

Answer all questions. Answers must be written within the answer boxes provided.

1. A cart X of mass 3.0 kg moving at 6.0 m s⁻¹ collides with a stationary cart Y of mass 6.0 kg. A spring is attached to X as shown.



The graph shows the velocities of X and Y before, during and after the collision.



The collision lasted for 40 ms.

•		٠	٠	•				٠											٠							٠	 ٠		٠				
	٠									. ,					 	. ,			,														
												. ,			 																		



(h)	lotormino	
(b)	Determine	
(/		

((i)	the magnitude of the average force exerted on Y,
((ii)	the average power delivered to Y,
((ii)	the average power delivered to Y,
((ii) 	the average power delivered to Y,
	(ii)	the average power delivered to Y,
	(ii)	

		(i	ii)	1	t	h	е	el	a	sti	ic	е	n	е	rg	У	S	tc	or	e	b	in	t	h	е	S	pr	in	ıg	а	ıt	t :	=	3	0 1	m	S.													2
							•								•			٠	•	•					٠							•		٠														 		
٠											٠																						٠				٠			٠	•		•		٠		٠			
																																							. ,									 		
																									•																									

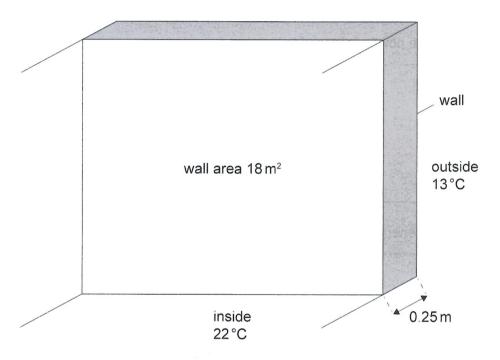


[2]

2.	(a)	Describe	the	mechanism	of	heat	transfer	hv	conduction	on

														 . ,										,			
					 				٠.					 											٠		

The diagram shows a wall separating the inside of a room from the outside. The temperature of the room is kept constant by a heater.



The following data are available:

Thickness of wall $= 0.25 \, \text{m}$

Area of wall = $18 \,\mathrm{m}^2$

Thermal conductivity of wall = 1.3 Wm⁻¹ K⁻¹

Constant room temperature = 22 °C

Constant outside temperature = 13 °C



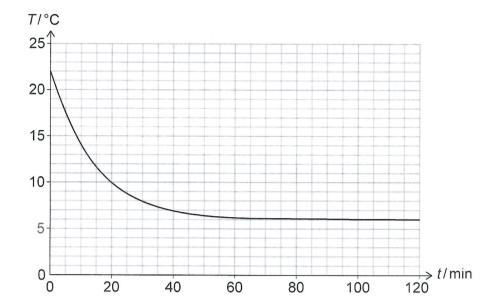
(k	0)														rm nit		31	er	ne	rg	Jy	IE	ea	ve	es	tr	1e	r	OC	om	l (l	nr	Οl	ug	n	เท	е	VV	aı	1 9	giv	VII	19	l	ne	<i>;</i>		[2]
	•		•	•		•	 	•				•																																				
				•								,		•		•	•							•					•																			
	•		*	•																				•											•					•			•					
	•		•							٠		•																																				



[3]

(Question 2 continued)

(c) At night the outside temperature falls below 13 °C. The heater is turned off at time t = 0. The graph shows the variation with time t of the temperature T of the room.



(i) Outline why the magnitude of the gradient of the graph decreases. [1]

••••••

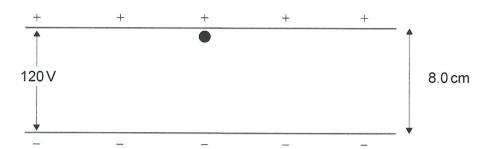
(ii) The pressure and volume of the air in the room remain constant but the number of molecules has increased. Determine the percentage increase in the number of molecules of air in the room between t = 0 and $t = 120 \,\text{min}$.



(0	(I)				do		е	Se	ec	0	nc	וג	a	N	O	Ι	ne	err	no	00	ıyı	18	ım	IIC	S	ap	р	IIE	es	tc) [ne	3 8	SIL	ua	ILI	or	1 (JΤ	ın	е	rc	00	m		[2	2]
																					. ,																										
										,																																					
																																. ,															
																																															- 1



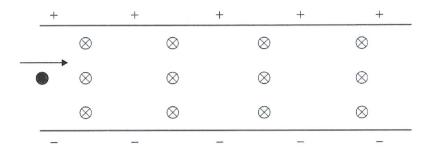
3. Two oppositely charged parallel plates are a distance 8.0 cm apart. The potential difference between the plates is 120 V. An alpha particle is placed on the positively charged plate and released from rest. Gravity is ignored.



- (a) Calculate the electric field between the plates. [1]
- (b) (i) Show that the acceleration of the alpha particle is about 7 × 10¹⁰ m s⁻². [2]
- (ii) Calculate the time taken for the alpha particle to reach the negative plate. [2]
- (iii) State, in eV, the kinetic energy of the alpha particle when it arrives at the negative plate. [1]



(c) A magnetic field directed into the plane of the page is now established between the plates. An alpha particle enters the region between the plates with a horizontal speed of $5.0 \times 10^5 \, \mathrm{m \, s^{-1}}$. The particle is not deflected.



Calculate	the	magnitude	of the	magnetic	field
Calculate	uie	magnitude	OI THE	magnetic	Held.

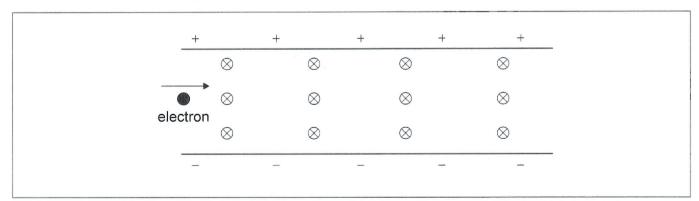
[2]

																								 	 	_	 -	 -		-	-
									٠	 	٠			•	 ٠	٠															
	٠			 ٠				 																				 ٠			

(d) The alpha particle in (c) is replaced by an electron. The electron enters the region between the plates with the same velocity as the alpha particle.

Draw, on the diagram, the path of the electron.

[1]



.....

4.	Two	sources of light produce an interference pattern on a screen.	
	(a)	Explain why the two sources need to be coherent for the interference pattern to be observed.	[2]
		• • • • • • • • • • • • • • • • • • • •	
	(b)	Light of wavelength 720 nm is incident on two narrow slits that are separated by 0.12 mm. An interference pattern is observed on a screen. P_1 and P_2 are the points of destructive interference closest to the central maximum M.	
		$\begin{array}{c c} & & & & \\ \hline \text{light} & & & & \\ \hline \end{array}$	
		P_2	
		screen	
		(i) Calculate, in radians, the angular separation φ of P_1 and P_2 .	[2]
		(ii) Suggest how energy conservation is consistent with the fact that the energy at P_1 and P_2 is zero.	[1]



A beam of light containing all wavelengths in the range [550 nm, 650 nm] is incident normally on a diffraction grating. The grating has 580 lines per mm. A diffraction pattern is observed on a screen as shown.



(c)	Outline why the interference at M is constructive.	[

(d)	Determine the largest order of the diffraction pattern in which all the wavelengths in the incident beam are present.	[3

(a)	Outline why Compton scattering offers better evidence for the particle nature of light than the photoelectric effect.	[2
	• • • • • • • • • • • • • • • • • • • •	
elect	oton of wavelength $6.40 \times 10^{-12} \mathrm{m}$ and energy $0.194 \mathrm{MeV}$ is incident on an tron at rest. The photon scatters at an angle θ with wavelength $7.47 \times 10^{-12} \mathrm{m}$. electron recoils.	
	scattered photon	
	θ	
	incident photon electron at rest	
	recoiling electron	
(b)	Outline why the wavelength of the scattered photon is longer than that of the incident photon.	[2
(c)	(i) Show that the energy of the scattered photon is about 0.17 MeV.	[1]

(This question continues on the following page)

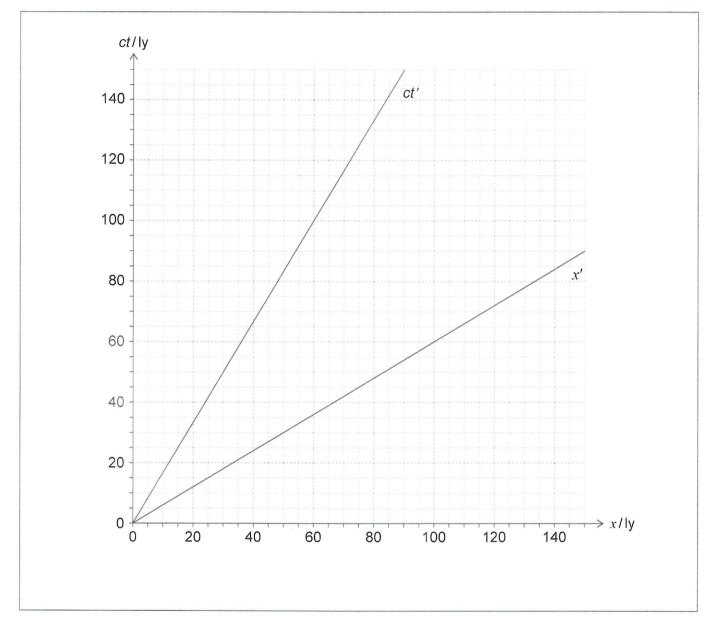


		(i	i)	(Ca	lc	ul	at	e	th	ne	k	in	e	tic	: 6	en	e	rg	y	0	ft	th	е	re	ec	oi	lir	าดู	1 6	ele	90	tr	or											[1]
					makes (man							-												-																					
	 				•														•										•			•		•			•				•				
•					•					٠															•				•			•					٠						•		

(iii) For an electron,
$$\frac{h}{m_e c} = 2.43 \times 10^{-12}$$
. Determine θ . [2]

												 											 										٠			
٠.												 										 	 													
												 										 									,					
												 										 												,		

6. Spacecraft S leaves Earth with speed 0.60c on the way to a planet P. As measured by Earth, P is at a distance of 30 ly from Earth. The diagram shows the spacetime axes of Earth (x, ct) and of S (x', ct').





(a)		rives at P after 50 years according to Earth. Calculate the time at which S arrives at cording to S clocks.	[2]
		•••••••••••••••••••••••••	
(b)	Whe	en S arrives at P, S sends a radio signal back to Earth.	
	(i)	Draw, on the spacetime diagram, the world line of the signal from when it is emitted until it is received on Earth.	[1]
	(ii)	Determine, using the spacetime diagram or otherwise, the time of travel of the signal according to S.	[3]



Turn over

7.	The following data are available for the Sun when it entered the main sequence.	
	$Mass = 2.0 \times 10^{30} kg$ $Radius = 7.0 \times 10^8 m$ $Surface \ temperature = 5800 K$ $Core \ temperature = 1.5 \times 10^7 K$ $Core \ density = 1.6 \times 10^5 kg m^{-3}$	
	(a) (i) State and explain which of the above features make nuclear fusion in the Sun possible.	[3]
	(ii) Outline how the Sun maintains its equilibrium state.	[2]



(b) The following sequence of nuclear fusion reactions takes place in the Sun.

Stage 1:
$${}^{1}_{1}H + {}^{1}_{1}H \rightarrow {}^{2}_{1}H + {}^{0}_{1}\overline{e}^{+} + {}^{0}_{0}v$$

Stage 2:
$${}_{1}^{1}H + {}_{1}^{2}H \rightarrow {}_{2}^{3}He + {}_{0}^{0}\gamma$$

Stage 3:
$${}_{2}^{3}$$
 He + ${}_{2}^{3}$ He $\rightarrow {}_{2}^{4}$ He + 2 $\times {}_{1}^{1}$ H

The nuclear mass of ${}^1_1\text{H}$ is 1.007276 u and that of ${}^2_1\text{H}$ is 2.013550 u. Estimate the energy released in stage 1 of these reactions.

[2]

* * * * * * *	 	

24ED17

(c)	ma: The	ss int e curr	o heli ent lu	um. T minos	he tot sity of t	al ener the Sur	uence v gy rele n is 3.8 s 1.5 ×	eased in $\times~10^{2}$	in the W. W	reaction	ons in	(b) is 4	3 × 10	⁻¹² J.	
	(i)	Sh	ow th	at the	Sun v	vill stay	on the	e main	sequ	ence f	or abo	ut 8 × 1	0° yea	rs.	[3]
	(ii)	Sta	ate or	ne ass	sumptio	on that	was m	iade in	gettir	ng the	answe	r to (c)	(i).		[1]
	an:	_													

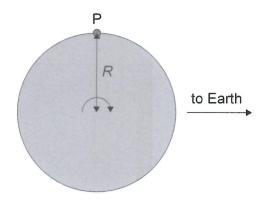
	 ([iii])	Es	stir	m	ate	е	th	e	tc	ota	al	m	ıa	SS	s I	os	st	by	y 1	th	е	Sı	ur	ı i	n	th	is	ti	m	e.											[2
																		۲.																									
																																								,	. ,		
															٠											٠																	

(0	d)	;	Tł Sı	ne ur	e s n's	su s s	rf sp	ac	ct	t ru	eı ın	m 1.	p	er	a	tu	re	9	of	t	h€	e ;	Sı	ur	ı i	s	58	80	00	K	΄.	D	et	er	m	nir	ne	tl	ne	p	е	ak	()	NE	IV	el	er	ng	th	il	٦.	th	е		[2
																																	-													-						-	-		



(i)

The Sun rotates about its axis. P is a point on the Sun's equator.



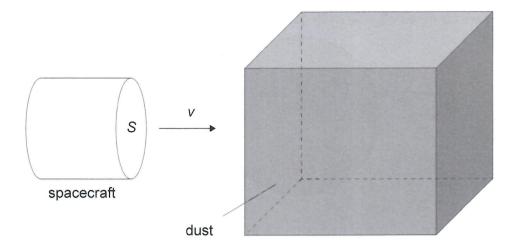
A particular spectral line of hydrogen from a laboratory source has wavelength 656.2797 nm. The same spectral line emitted from P has wavelength 656.2753 nm when measured on Earth.

(i)	I	Ξx	ola	iin	th	is	ol	os	er	Vä	ati	io	n.																																					[2]
 																																																	,			
 (ii)		Cal	cu	ıla	te	th	е	pe	erio	00	d (of —		-		lu	ıti	or	ו (—	of	tł	ne	. 5	Su	ın																									[3]]
 		,																																													,					
 												٠																	٠							٠			٠				٠									
 															,																																٠					
	((ii) ((ii) Cal	(ii) Calcu	(ii) Calcula	(ii) Calculate	(ii) Calculate th	(ii) Calculate the	(ii) Calculate the pe	(ii) Calculate the peri	(ii) Calculate the period	(ii) Calculate the period of	(ii) Calculate the period of	(ii) Calculate the period of re	(ii) Calculate the period of rev	(ii) Calculate the period of revo	(ii) Calculate the period of revolu	(ii) Calculate the period of revolution of	(ii) Calculate the period of revolution of the	(ii) Calculate the period of revolution of the	(ii) Calculate the period of revolution of the S	(ii) Calculate the period of revolution of the Su	(ii) Calculate the period of revolution of the Sun	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.	(ii) Calculate the period of revolution of the Sun.			



Turn over

8. (a) A cylindrical spacecraft of cross-sectional area S moves with velocity v. The spacecraft enters a region of dust of density ρ . All the dust that comes into contact with the forward cross-sectional area of the spacecraft sticks to the spacecraft increasing its mass.



(i) Show that in time Δt the mass of the spacecraft will increase by an amount $\rho Sv\Delta t$. [1]

(ii)) Deduce that the	spacecraft will experience a dr	ag force $ ho S v^2$ opposite to its velocity.	[2



(b) A satellite of mass m is in a circular orbit of radius r around Earth. Earth has mass M.

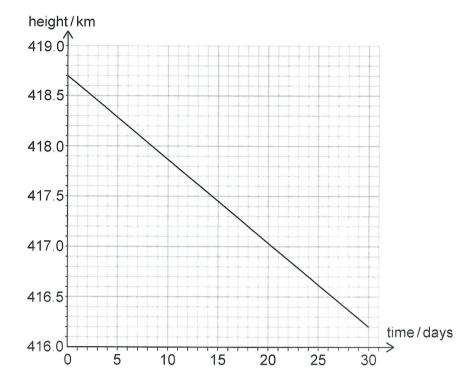
(i) Show that the total energy of the satellite is given by $E_{T} = -\frac{GMm}{2r}$. [2]

(ii) The satellite experiences a drag force due to the atmosphere of Earth. With reference to the results in (a) and (b)(i), state and explain the likely fate of this satellite.

[4]

		 					 										 							. ,				
•				٠	 																					. ,		

(c) The graph shows the variation with time of the height of the International Space Station (ISS) during a 30-day period.



The following data are available:

Mass of Earth = 5.98×10^{24} kg Mass of the ISS = 4.20×10^{5} kg Radius of the Earth = 6.38×10^{6} m

For the 30-day period indicated in the graph the loss of total energy of the ISS is $4.5\times10^9\,J_{\cdot}$

(i) Calculate the average rate at which the energy of the ISS is being dissipated. [2]



(ii)	Show, using the average height during the 30-day period, that the speed of the ISS is about $8\times10^3\text{m}\text{s}^{-1}$.	[2]
(iii)	Estimate the increase in the temperature of the ISS assuming all the lost energy went into thermal energy of the ISS. Take the specific heat capacity of the ISS to be $500\mathrm{Jkg^{-1}K^{-1}}$.	[2]
(iv)	Estimate, using the answers to (c)(i) and (c)(ii), the average drag force that acted on the ISS.	[3]



Turn over

(d)	At the end of the 30-day period, rockets are fired to bring the ISS back to its initial
,	height. The energy density of liquid hydrogen rocket fuel is $8.5 \times 10^3 \text{MJ} \text{m}^{-3}$.

Estimate the volume of fuel needed.										

